

Simulation: A Bluffer's Guide.

A brief journey into the rôle, form and limitations of computer simulation in condensed matter physics.

(distinctly non-exhaustive, rather subjective and not a lot of quantum mechanics either.)

An attempt to place my work into context.



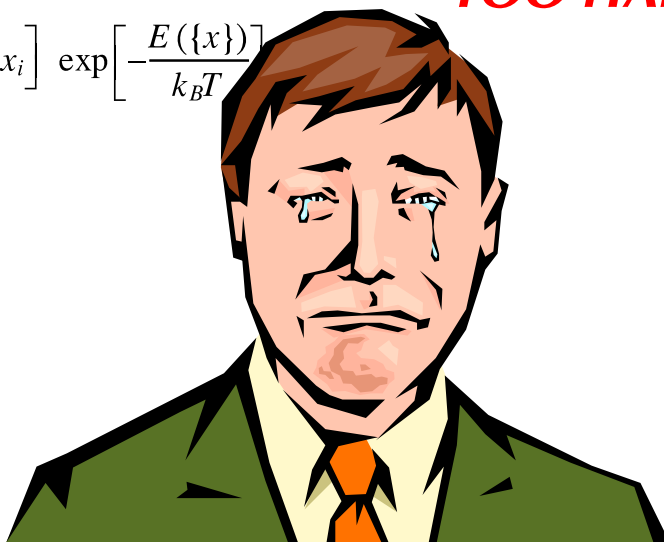
Also, an excuse for using clip-art gratuitously.

Why Simulate?

$$E(\{x\}) = \sum_{i=1}^N \sum_{j=i+1}^N V(r_{ij})$$

**THE THEORY'S JUST
TOO HARD!**

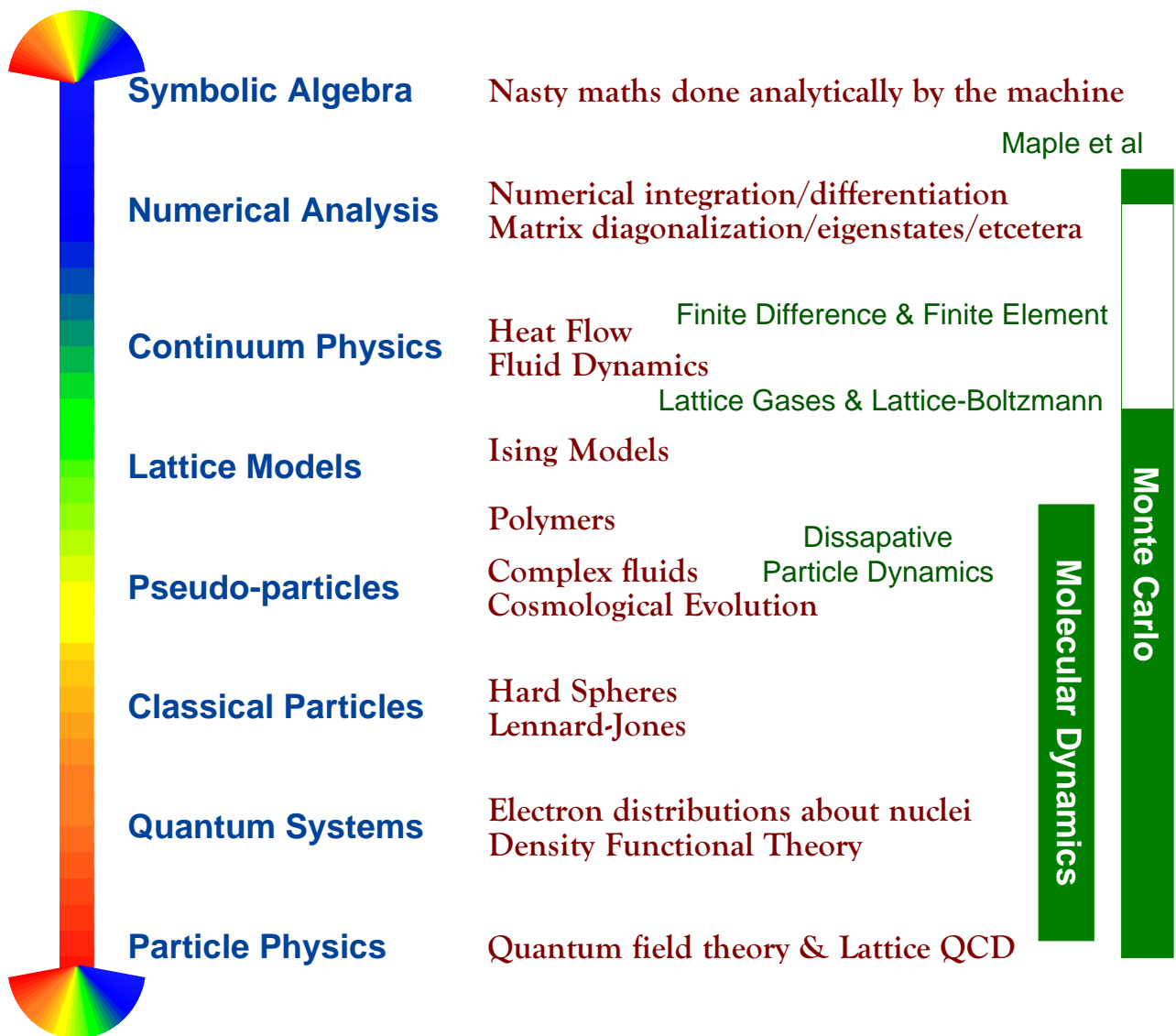
$$Z(N, V, T) = \prod_{i=1}^{3N} \left[\int_{-\infty}^{+\infty} d\dot{x}_i \right] \prod_{i=1}^{3N} \left[\int_{-\infty}^{+\infty} dx_i \right] \exp \left[-\frac{E(\{x\})}{k_B T} \right]$$



Or if you want a second opinion, why not ask your friendly neighbourhood coder.

~~The~~ A Spectrum of Computational Physics

Most definitely both arbitrary and non-exhaustive.



Overall Aims

Given some model many-body system.
(A finite set of objects in some space/geometry
with some rules for how they interact)

We wish to predict observable properties of a real system.
(Eg bulk properties an effectively infinite number of objects).

Equilibrium Properties

Free-energies, phase behaviour, compressibility, specific heat...

For $T = 0K$, calculating the energy is enough.

For $T > 0K$, we need lots of microstates.

Dynamic Properties

Relaxation times, diffusion times...

Needs lots of microstates.



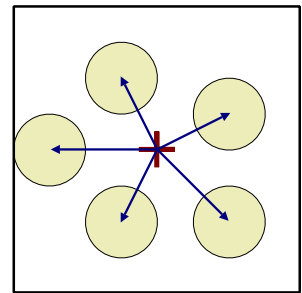
The Simulation Arena

Characterizing the System

+1	+1	+1	+1
-1	-1	+1	+1
-1	-1	+1	-1
-1	-1	-1	+1

Ising Model

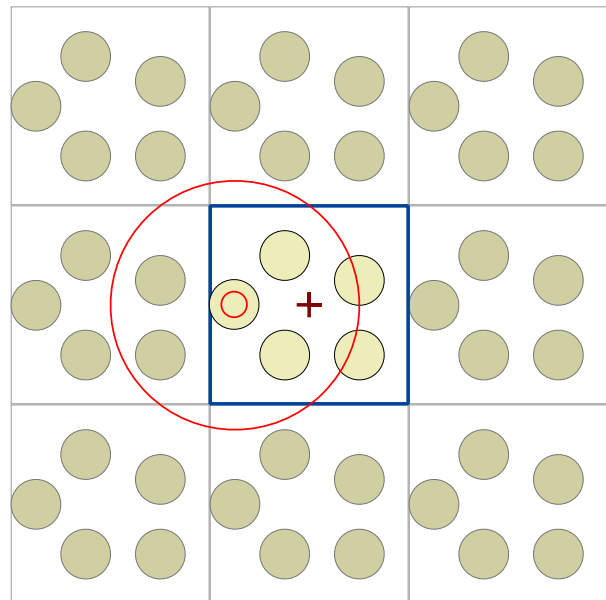
Hard Spheres



+1	+1	+1	+1	-1	-1	+1	+1	-1	-1	+1	-1	-1	-1	-1	+1
----	----	----	----	----	----	----	----	----	----	----	----	----	----	----	----

Periodic boundary conditions

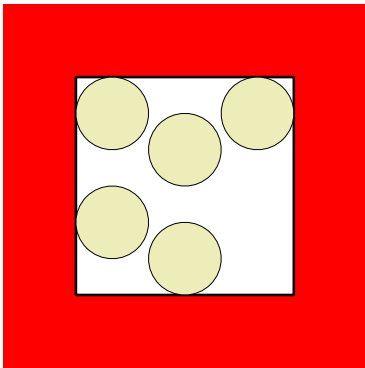
+1	+1	+1	+1	+1	+1	+1	+1	+1	+1	+1	+1
-1	-1	+1	+1	-1	-1	+1	+1	-1	-1	+1	+1
-1	-1	+1	-1	-1	-1	+1	-1	-1	-1	+1	-1
-1	-1	-1	+1	-1	-1	-1	+1	-1	-1	-1	+1
+1	+1	+1	+1	+1	+1	+1	+1	+1	+1	+1	+1
-1	-1	+1	+1	-1	-1	+1	+1	-1	-1	+1	+1
-1	-1	+1	-1	-1	-1	+1	-1	-1	-1	+1	-1
-1	-1	-1	+1	-1	-1	-1	+1	-1	-1	-1	+1
+1	+1	+1	+1	+1	+1	+1	+1	+1	+1	+1	+1
-1	-1	+1	+1	-1	-1	+1	+1	-1	-1	+1	+1
-1	-1	+1	-1	-1	-1	+1	-1	-1	-1	+1	-1
-1	-1	-1	+1	-1	-1	-1	+1	-1	-1	-1	+1



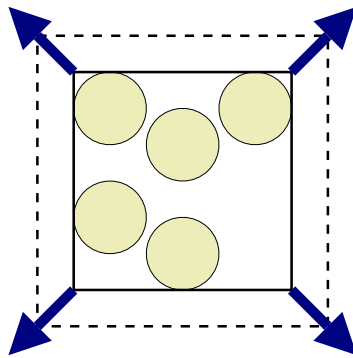
External influences

Microcanonical ensemble OR

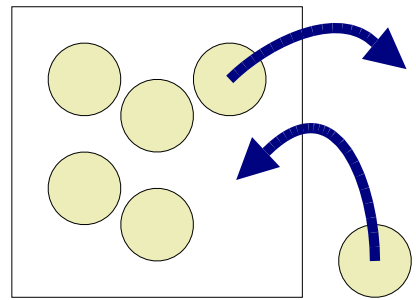
Fix temperature



Fix pressure



Fix chemical potential



Using appropriate statistical mechanics rules.

Dimensionless Units

Scale out “real” units, and capture the physics.

Hard Spheres:

Given N , only the ratio between radius and volume matters.

r^3/V : ie the density, ρ^* .

Ising Model:

The ratio of the interaction energy to the thermal energy matters:

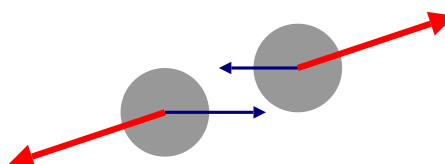
$J/k_B T = 1/T^*$.

Molecular Dynamics

Eg Evolve positions of particles under Newton's Laws.

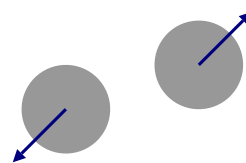
$$a = \frac{dv}{dt} \approx \frac{\Delta v}{\Delta t} \quad v = \frac{dx}{dt} \approx \frac{\Delta x}{\Delta t}$$

1) Calculate the forces:



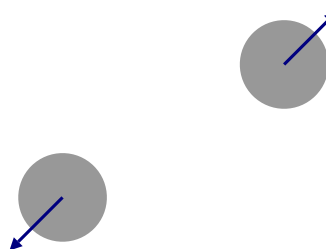
2) Update the velocities:

$$v(t + \Delta t) = v(t) + a \Delta t$$



3) Update the positions:

$$x(t + \Delta t) = x(t) + v \Delta t$$



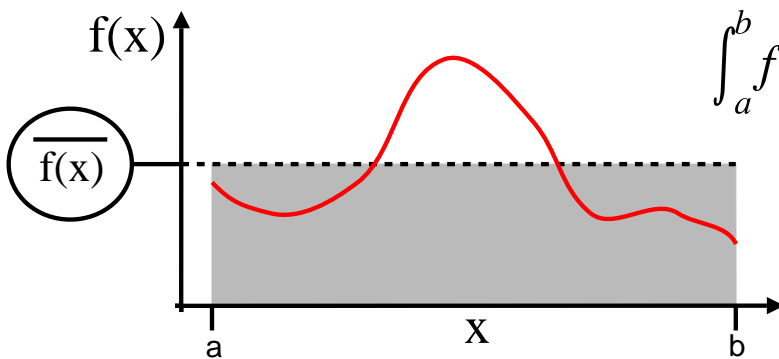
4) Repeat 1, 2 and 3.

There are more accurate ways!

Getting fixed temperature/pressure etc requires mild bludgeoning.

Monte Carlo Method

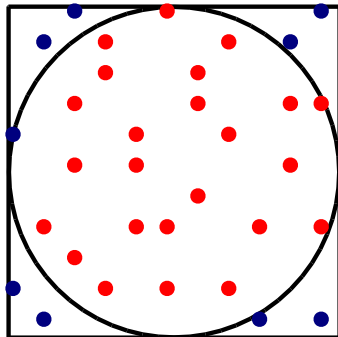
In 1 Dimension:



$$\int_a^b f(x) dx \equiv (b - a) \overline{f(x)}$$

$$\approx \frac{1}{n} \sum_{i=1}^N f(\zeta_i)$$

In Two Dimensions:



$$\frac{A(\text{circle})}{A(\text{square})} = \frac{\pi r^2}{(2r)^2} = \frac{\pi}{4} \approx \frac{n_{\text{inside}}}{n_{\text{total}}}$$

$$\left[\text{e.g. } \frac{24}{33} = 2.9 \right]$$

Can we do this for many-dimension integrations over phase space?

$$Z(N, V, T) = \prod_{i=1}^{3N} \left[\int_{-\infty}^{+\infty} d\dot{x}_i \right] \prod_{i=1}^{3N} \left[\int_{-\infty}^{+\infty} dx_i \right] \exp \left[-\frac{E(\{x\})}{k_B T} \right]$$

Monte Carlo In Statistical Mechanics:

Eg In the canonical ensemble.

Generate microstates at random.

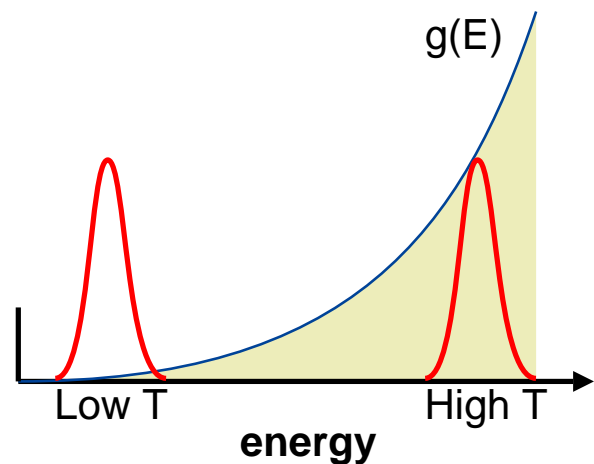
Estimate energy using...

$$\bar{E} \approx \frac{1}{n} \sum_{i=1}^n E_i \exp[-\beta E_i]$$

This doesn't work!

Eg Ising Model:

Tend to generate 'hot' configurations.

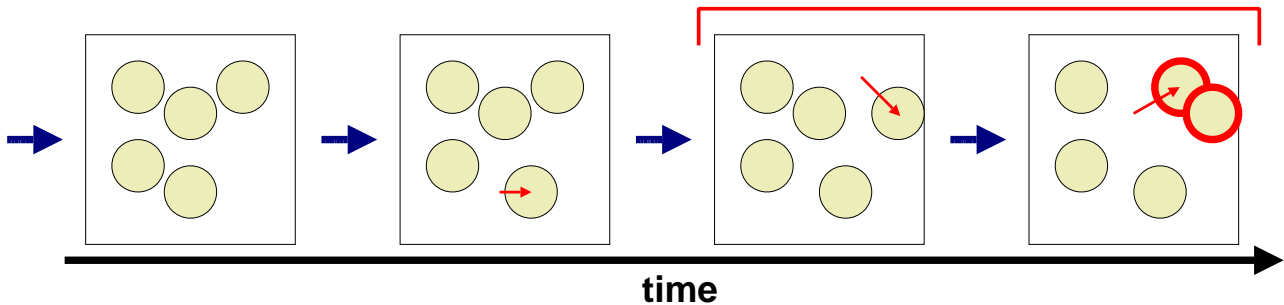


Eg Hard-Spheres:

Doesn't work at high densities.

Can we generate the right (high Boltzmann weight) configurations all the time, instead of waiting for them?

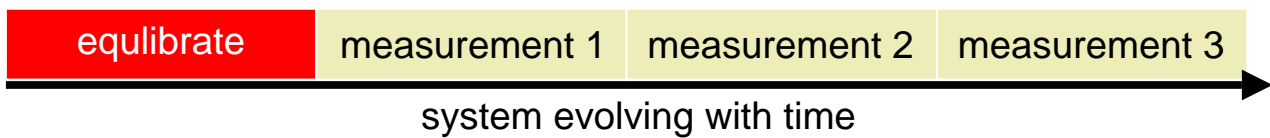
Markov Chains:



OK, but we can't accept all moves.

How do we decide whether to accept a trial move?

Convergence:



Can show that **microscopic reversibility** ('detailed balance') is a **sufficient** condition to ensure convergence.

$$P(i) P(i \rightarrow j) = P(j) P(j \rightarrow i)$$

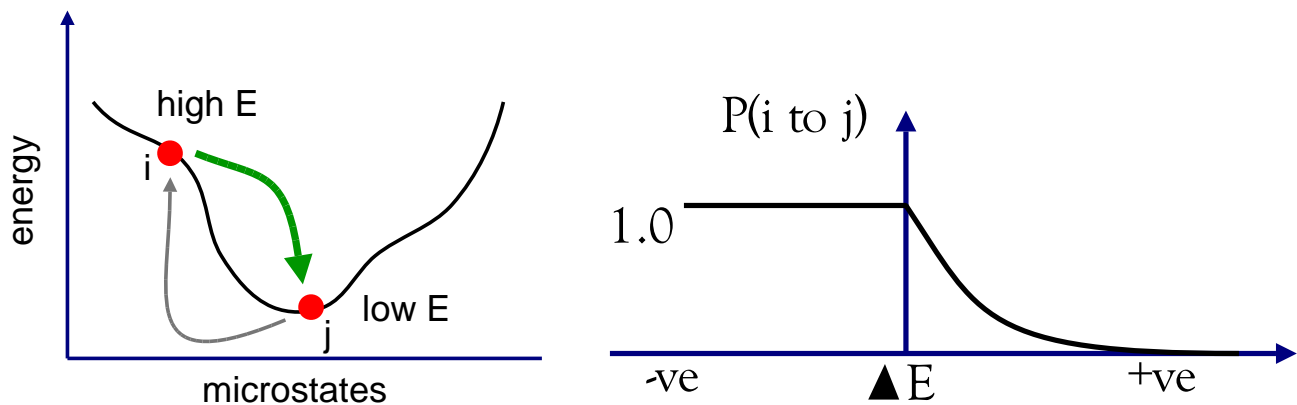
We need an algorithm that satisfies this condition and will sample with the Boltzmann distribution.



Importance Sampling:

Moves to lower energies are **ALWAYS** accepted.

Moves up are accepted **sometimes...**



$$P(i \rightarrow j) = \min[1.0, \exp(-\beta(E_j - E_i))]$$

$$P(i) = \frac{1}{Z} \exp(-\beta E_i)$$

Putting these into the detailed balance condition:

$$\exp(-\beta E_i) \times 1.0 = \exp(-\beta E_j) \times \exp(-\beta(E_i - E_j))$$



This generates a sequence of Boltzmann distributed microstates.

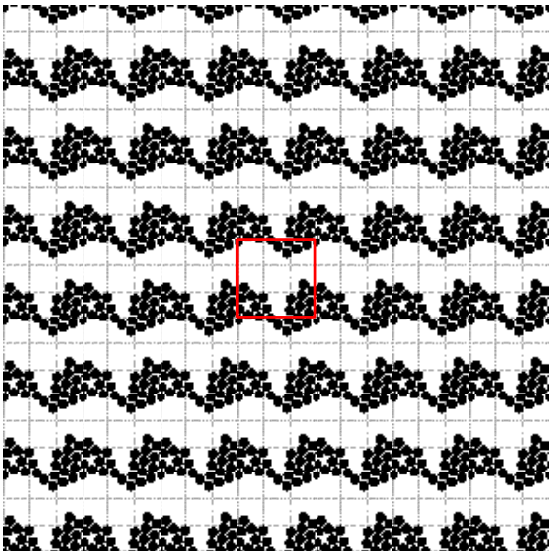
Estimate observables like this:

$$\bar{E} \approx \frac{1}{n} \sum_{i=1}^n E_i$$

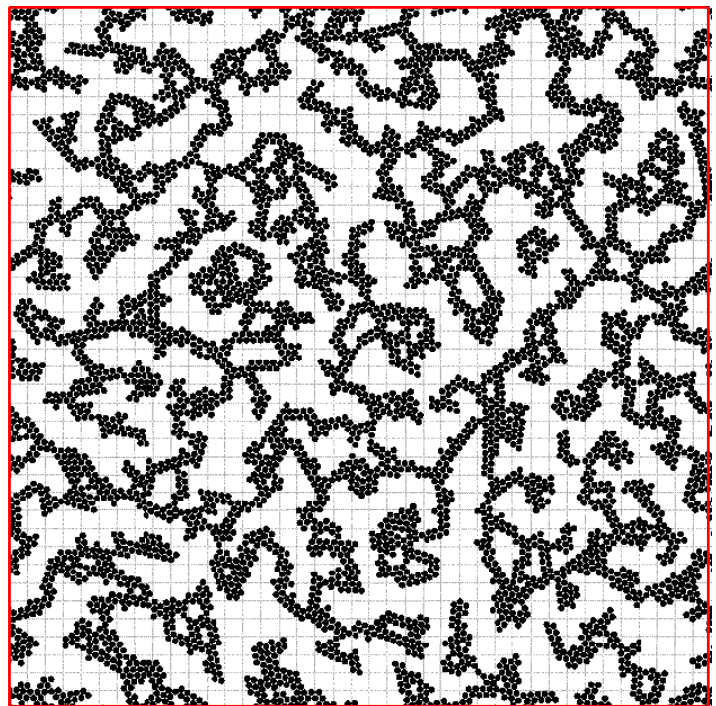
Limitations: *Finite System Size*

Example: 2D sticky spheres:

$N = 25$



$N = 5000$



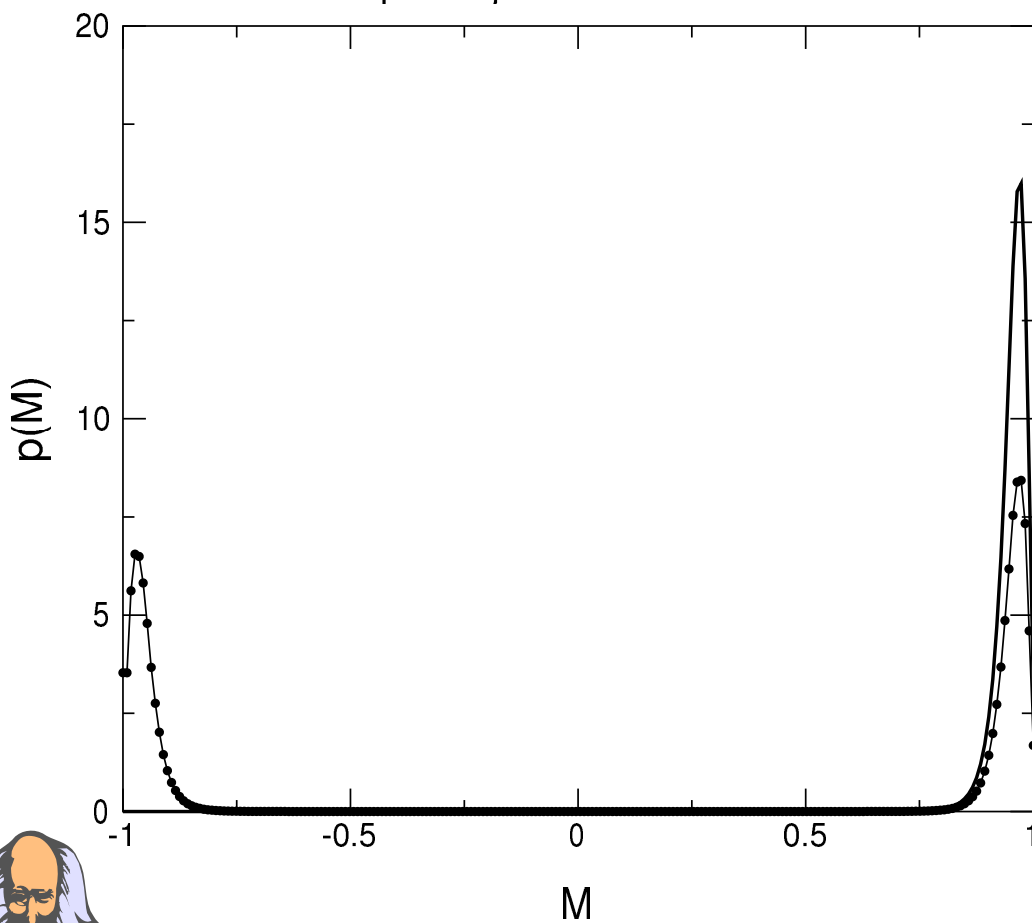
One should always examine a range of system sizes.

A wealth of literature exists on finite size effects and theory to cope with them ('finite size scaling theory').

Limitations: *Finite Simulation Time*

Example: Ising model at low temperature:

15x15 spins -- $\beta = 0.55$ -- 1.0×10^6 MCS

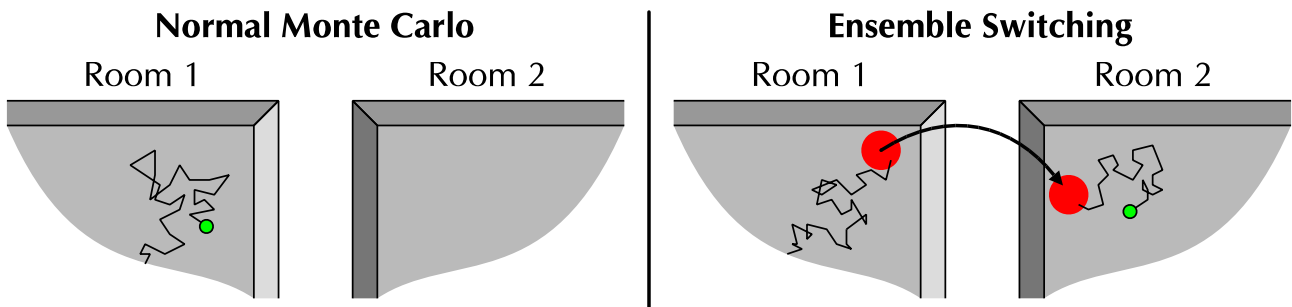


With MC is that you can fudge things **legally**.

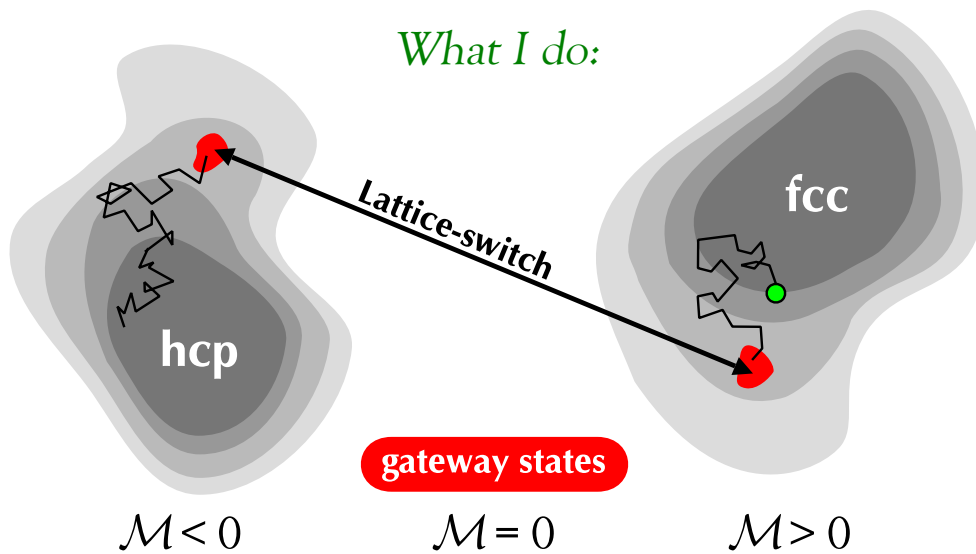


Beating The System

Changing Rooms



What I do:



Final Remarks...

Notable Omissions:

Stability of algorithms.

Finite precision problems.

*Estimating accuracy under Monte Carlo,
(and about a million other details).*

Quantum Mechanics:

DFT makes calculations 10-100 times harder.

Supercomputing:

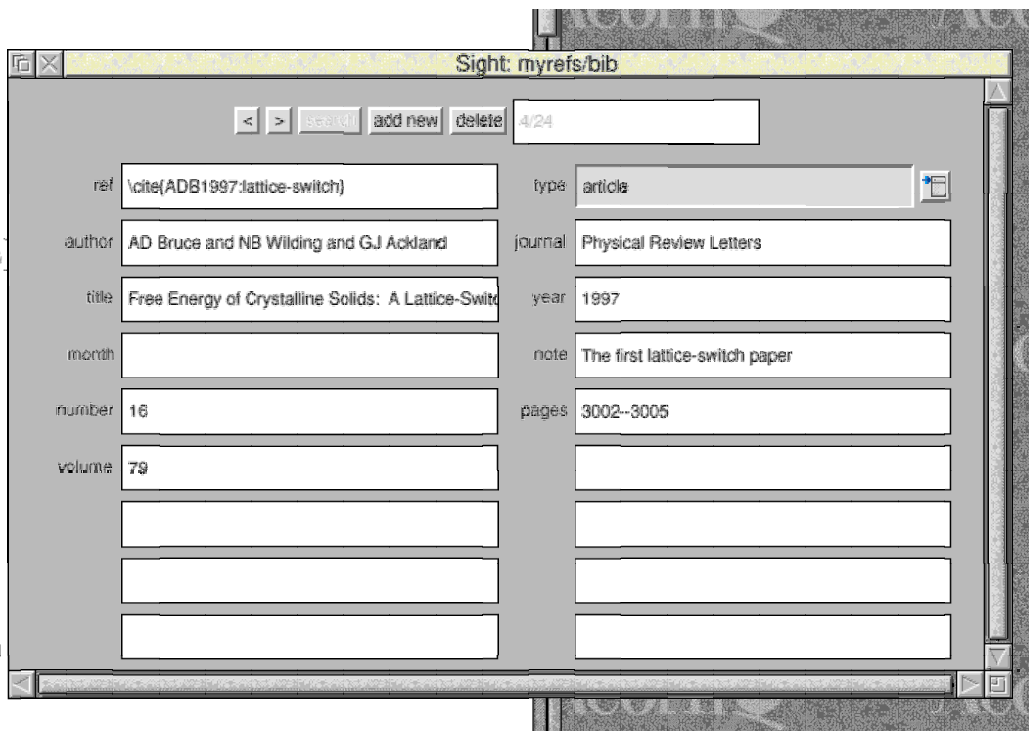
Various tricks and cunning techniques.

Makes simulation 10-100 faster.

And hopefully you don't
feel like this.



Sight: BibTeX Citation Management Program



$E(x) =$
 $Z(N, V, T) = \prod_{i=1}^N \int_{-\infty}^{\infty} dx_i$

ref	<code>\cite{ADB1997:lattice-switch}</code>	type	article
author	AD Bruce and NB Wilding and GJ Ackland	journal	Physical Review Letters
title	Free Energy of Crystalline Solids: A Lattice-Switch	year	1997
month		note	The first lattice-switch paper
number	16	pages	3002-3005
volume	79		

Allows you to edit, add and delete entries in your BibTeX citation file in a "nice" way.

On Unix (at least), it makes it very easy to insert the right `\cite{}`.

Written in Java - can run on any kind of machine.